

# Basic structure of Monte Python

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- 1 Code Structure
- 2 Important Modules in (some) detail
- 3 Usage
- 4 Practice with Monte Python

# Outline

## 1 Code Structure

- Wrapper
- Flow

## 2 Important Modules in (some) detail

## 3 Usage

## 4 Practice with Monte Python

# classy - wrapper around CLASS

## Cython

- language to **interface** Python with C
- can be used to **speed up** a bottleneck in Python or...
- to wrap an **existing C code** and call it from Python as a normal function

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## Classy

- `classy.pyx` **wraps** CLASS modules and some functions (only the ones needed from within Monte Python).
- uses a **dictionary** in place of the **.ini file** to give parameters (file content)
- if no arguments are specified, CLASS **default values** will be used.

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## Classy

- `classy.pyx` **wraps** CLASS modules and some functions (only the ones needed from within Monte Python).
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- if no arguments are specified, CLASS **default values** will be used.

## Last word

You (most probably) don't need to know how `classy.pyx` is written. The only important thing is **its interface**.

# General structure of the modules

## Files

### Input Files

- configuration file: **path to codes on machine**
- parameter file: **parameters, prior range, proposal, experiments**

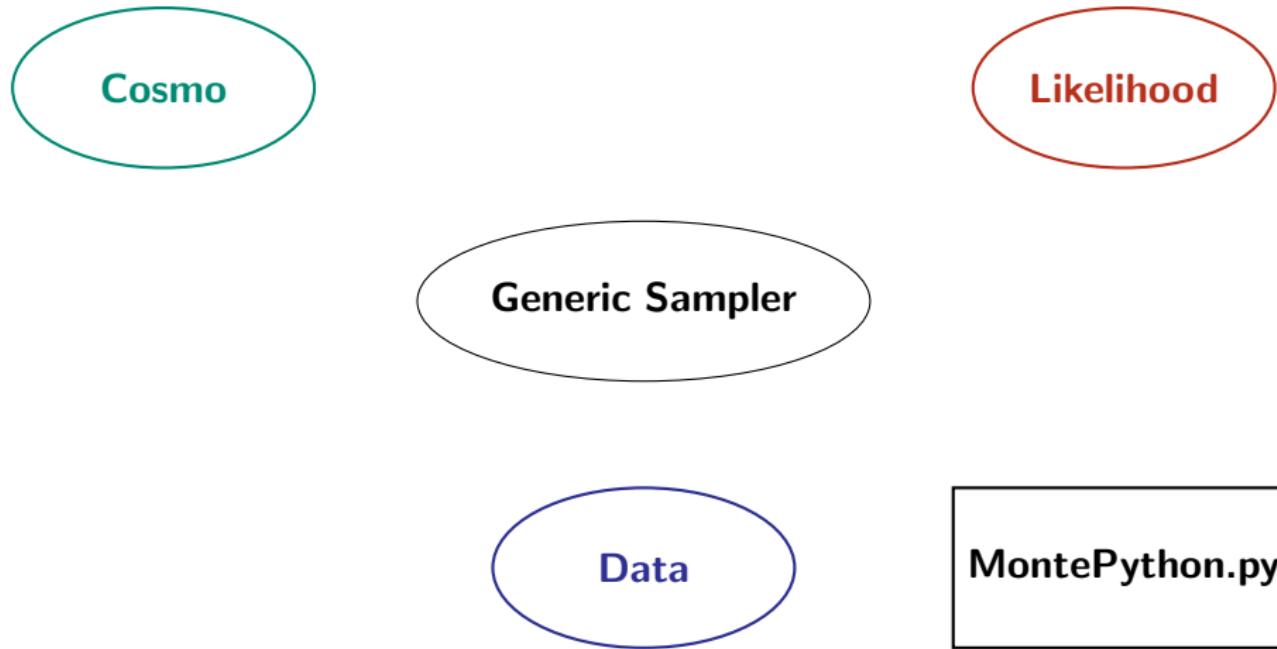
opt covariance matrix, bestfit file

### Output

- a folder: **stores every information concerning the run**
- a chain per run: **Markov Chain of a given length**

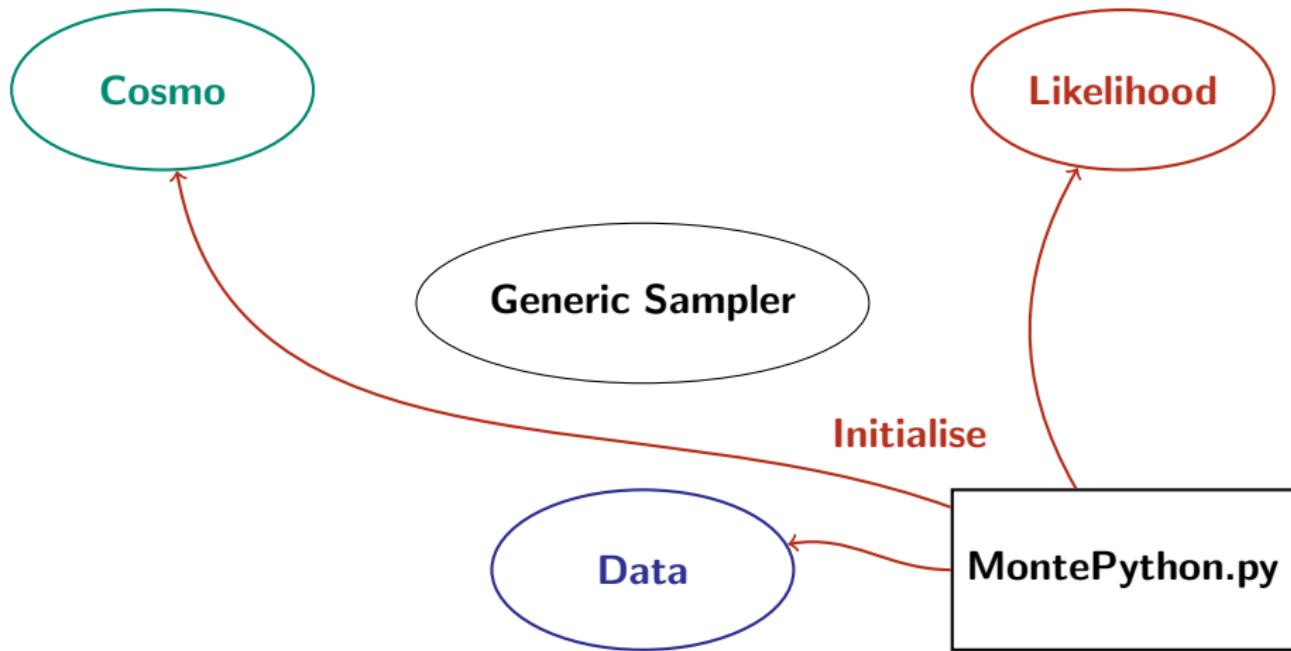
# General structure of the modules

A drawing



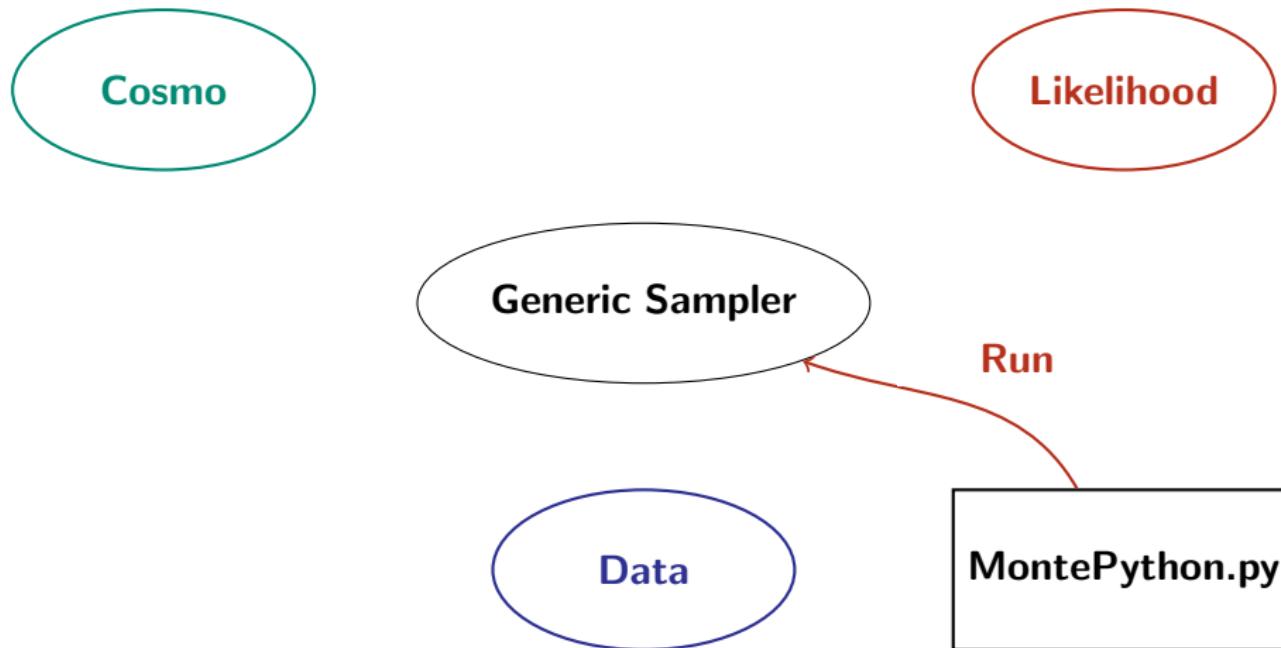
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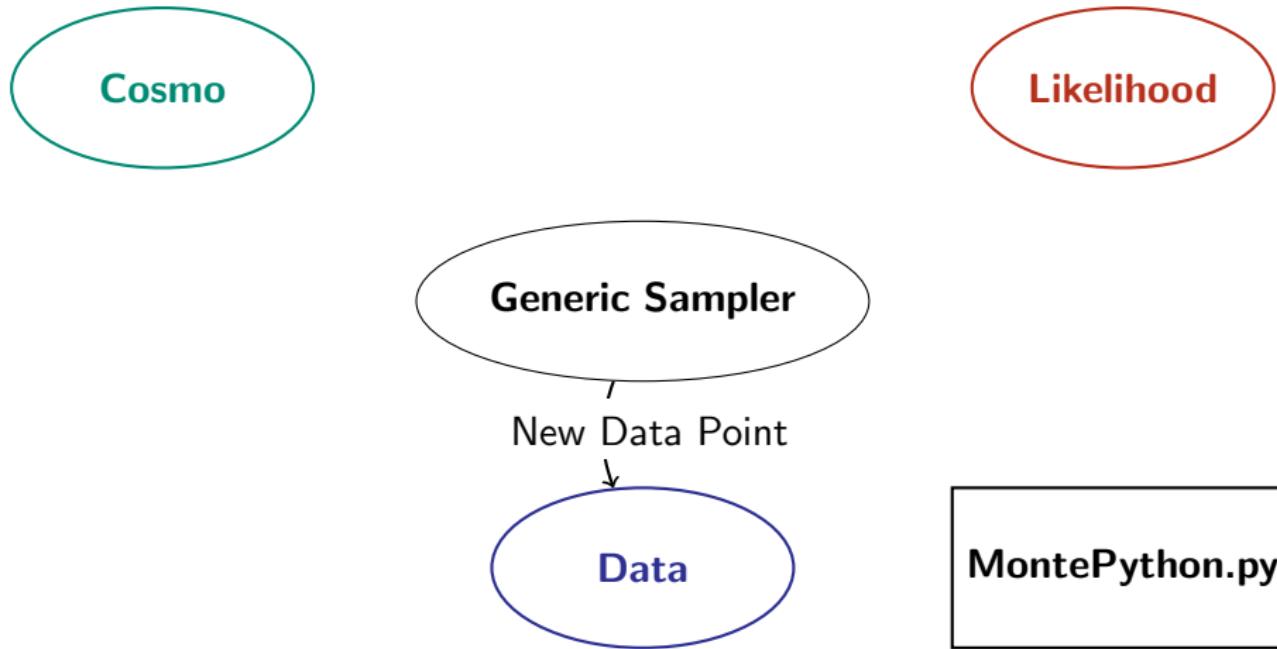
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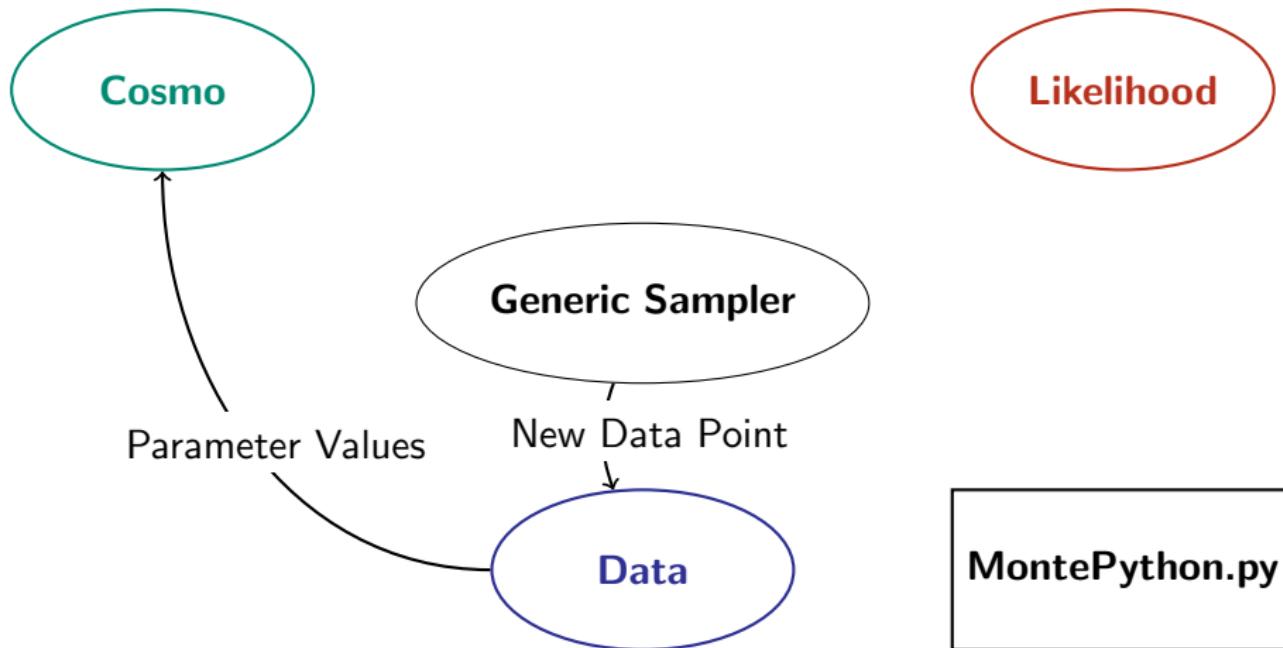
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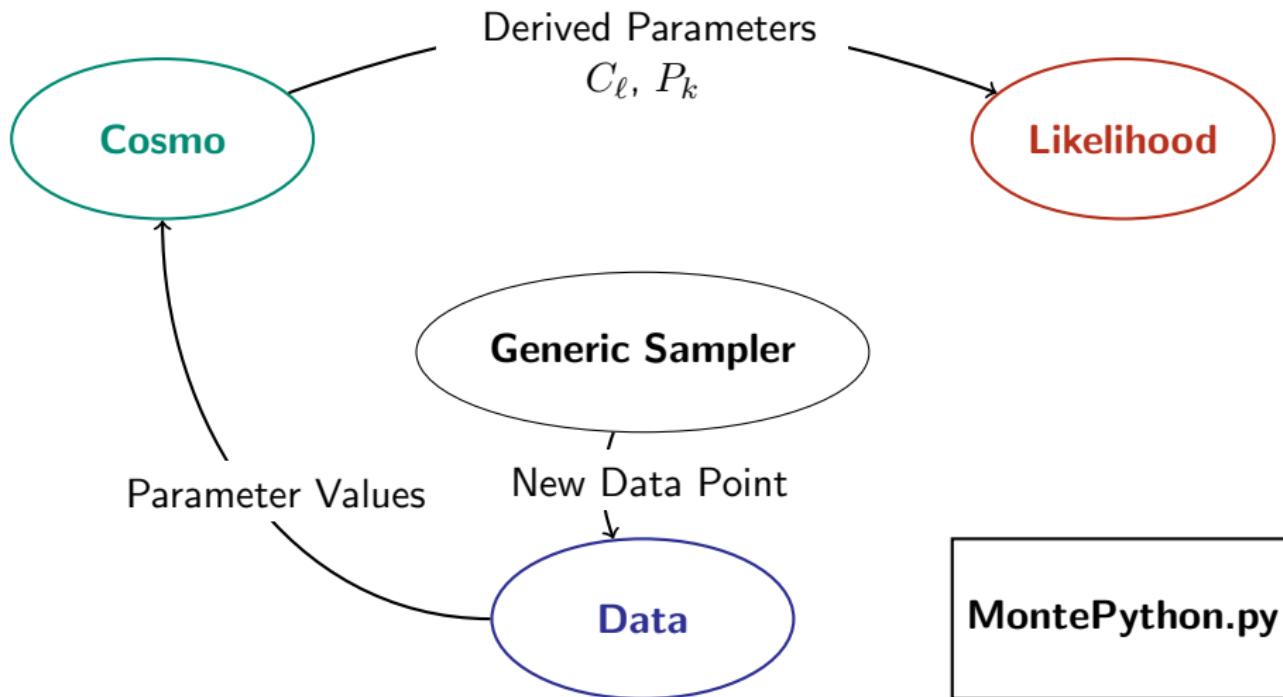
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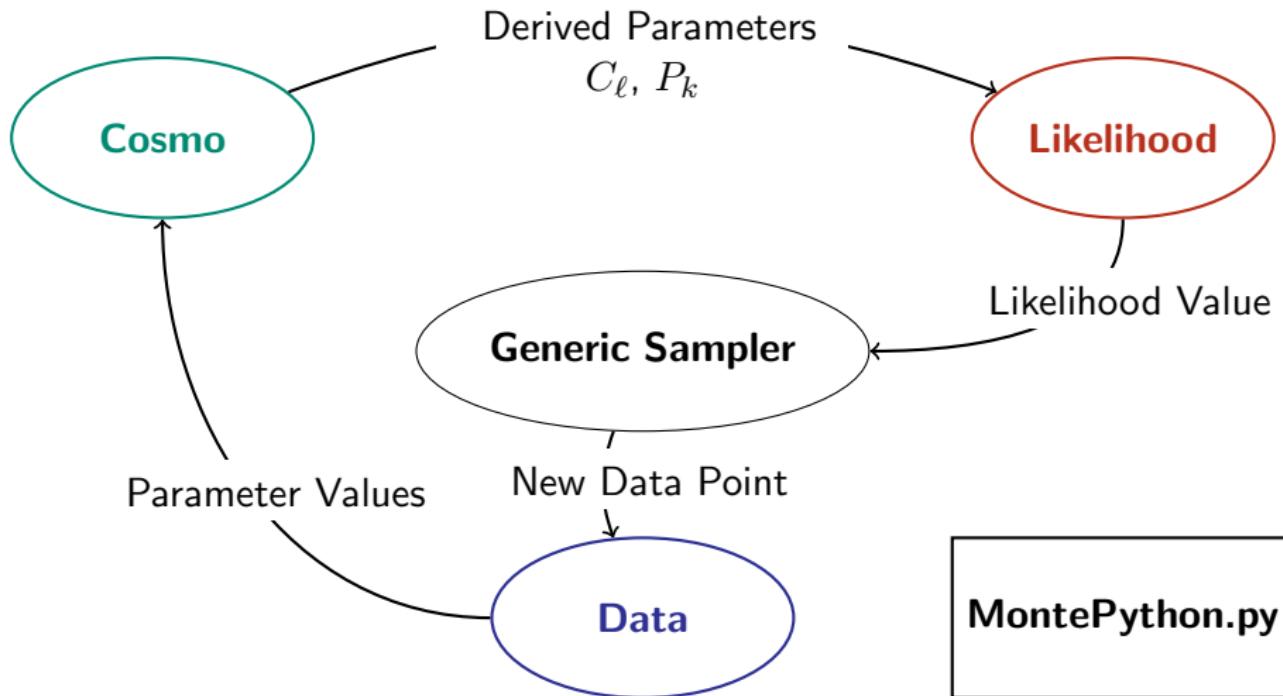
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# Information containers

Foreword: class definitions in capital letters, instances in small.

## Classes

- **Data** defined in `data.py`
- **Class** defined in `classy.pyx`
- **Likelihood** defined in `likelihood_class.py`

# Information containers

Foreword: class definitions in capital letters, instances in small.

## Classes

- **Data** defined in `data.py`
- **Class** defined in `classy.pyx`
- **Likelihood** defined in `likelihood_class.py`

## instances

- **data** initialized in `initialise.py`
- **cosmo** initialized in `initialise.py`
- **hst, bicep2, ...** initialized in `initialise.py`

# General structure of the modules

## Main Modules

- MontePython **Simple script launching the code**

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- **run call initialise, and launch a sampler session**

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- **parser\_mp** **reads the command line arguments**

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## Helper Modules

- `analyze` Computes convergence, posterior from chains
- `io_mp` Handles I/O, error message

# Outline

1 Code Structure

2 Important Modules in (some) detail

- MontePython.py
- Initialise.py
- data.py
- sampler.py
- likelihood class
- analyze.py

3 Usage

4 Practice with Monte Python

# Note on the documentation

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## Use it!

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Go to the issues page

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## Use the wiki!

[https://github.com/baudren/montepython\\_public/wiki](https://github.com/baudren/montepython_public/wiki), or  
[https://github.com/lesgourg/class\\_public/wiki](https://github.com/lesgourg/class_public/wiki) for classy business.

# MontePython.py

## Role

Convenience script that calls the Monte Python run function.

# Initialise I

## Main

- Reads command line, configuration file

```
29 # Parsing line argument
30 command_line = parser_mp.parse(custom_command)
31
32 # Recovering the local configuration
33 path = recover_local_path(command_line)
```

# Initialise I

## Main

- Reads command line, configuration file
- **Creates a data instance**

56

```
data = Data(command_line, path)
```

# Initialise I

## Main

- Reads command line, configuration file
- Creates a data instance
- **Initializes the cosmological module**

```
72     # Loading up the cosmological backbone. For the moment, only
    CLASS has been
73     # wrapped.
74     cosmo = recover_cosmological_module(data)
```

# Data I

## Defining a data class

- Initialization

```
35 class Data(object):
36     """
37     Store all relevant data to communicate between the different
38     modules.
39     """
40
41     def __init__(self, command_line, path):
42
43         self.cosmo_arguments = {}
44
45         self.mcmc_parameters = od()
```

# Data I

## Defining a data class

- Initialization
- **Fill in parameter information**

```
195      # Read from the parameter file to fill properly the
196          mcmc_parameters
197      # dictionary.
198      self.fill_mcmc_parameters()
```

# Data I

## Defining a data class

- Initialization
- Fill in parameter information
- **Log parameter file if needed**

# Data II

## Defining a data class

### • Initialization of likelihood (dynamical)

```
338     for elem in self.experiments:  
  
343         # ... import easily the likelihood.py program  
344         exec "from likelihoods.%s import %s" % (  
345             elem, elem)  
  
350         exec "self.lkl['%s'] = %s('%s/%s.data',\br/>351             self, command_line)" % (  
352                 elem, elem, folder, elem)
```

# Data II

## Defining a data class

- Initialization of likelihood (dynamical)

## Why so complicated

No hard coded likelihood! The code does not know the names: **no need to modify the core code to add a new likelihood**

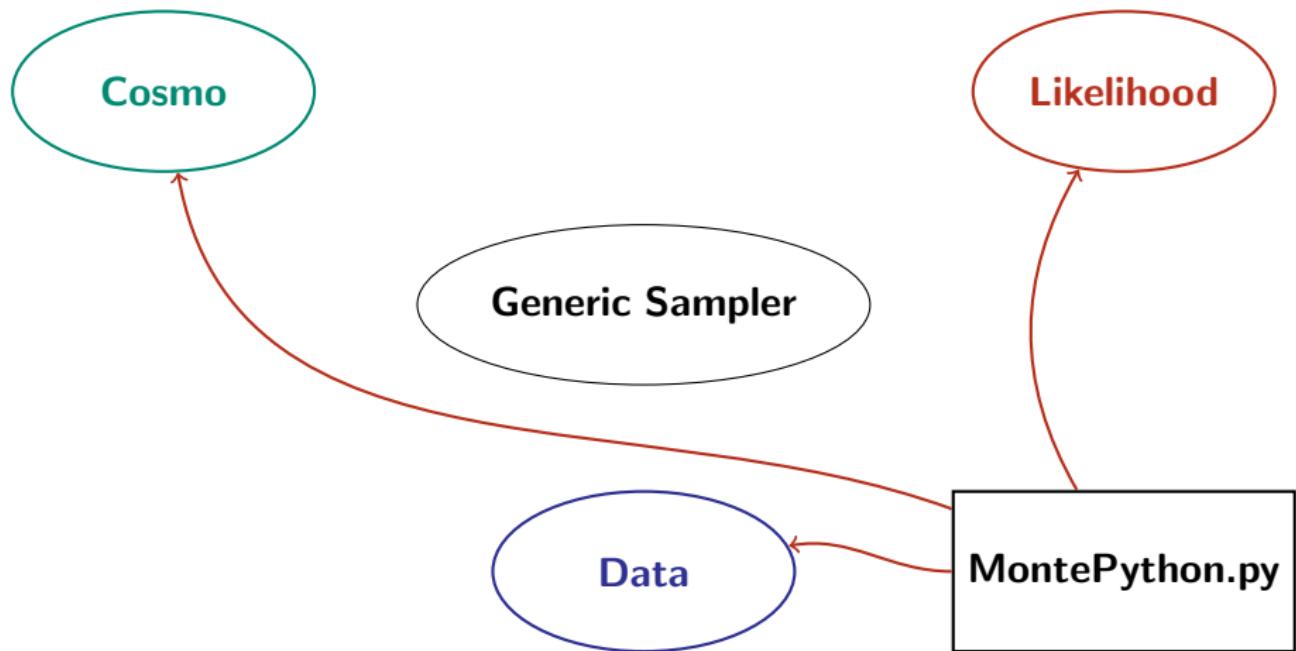
# Data III

## essential function

`get_mcmc_parameters` returns the list of desired parameters.

- `get_mcmc_parameters(['varying'])`
- `get_mcmc_parameters(['cosmo', 'nuisance'])`
- `get_mcmc_parameters(['cosmo', 'varying'])`

# Recap Initialisation



# Sampler I

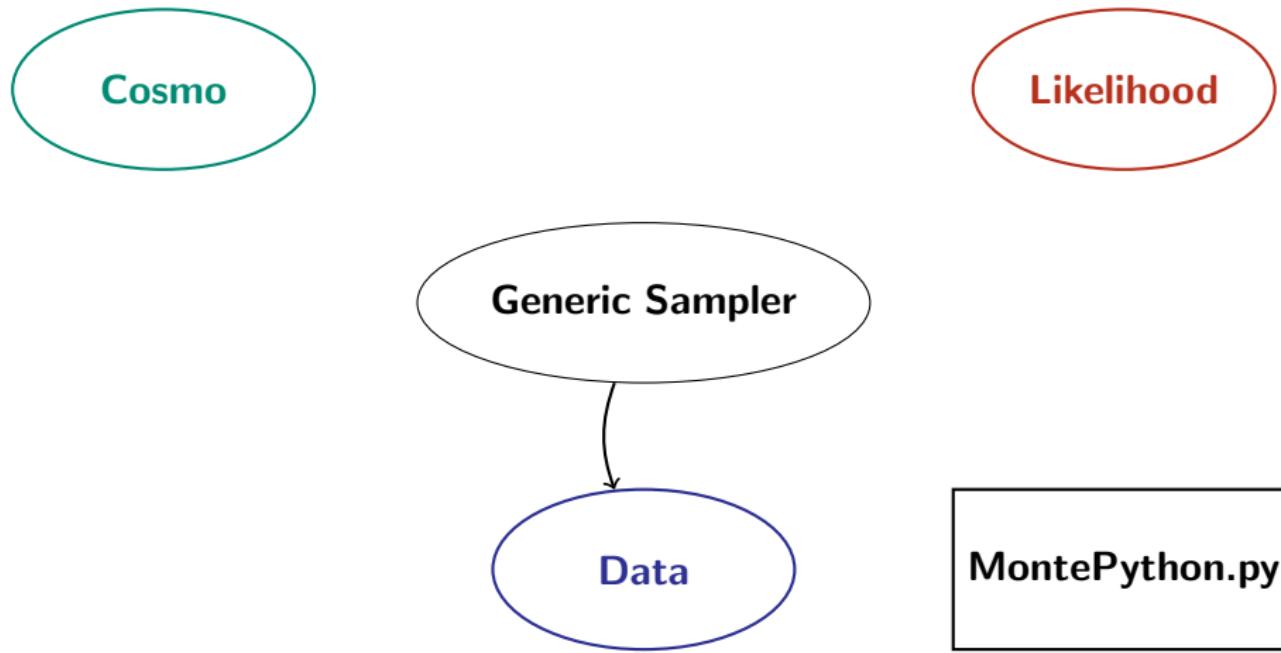
## Generic helper functions

- `compute_llkl(cosmo, data)`
- `get_covariance_matrix(data, command_line)`

## Role

calling the sampler specified via the **command line**

# Choosing a new point



# Sampler II

Get new position

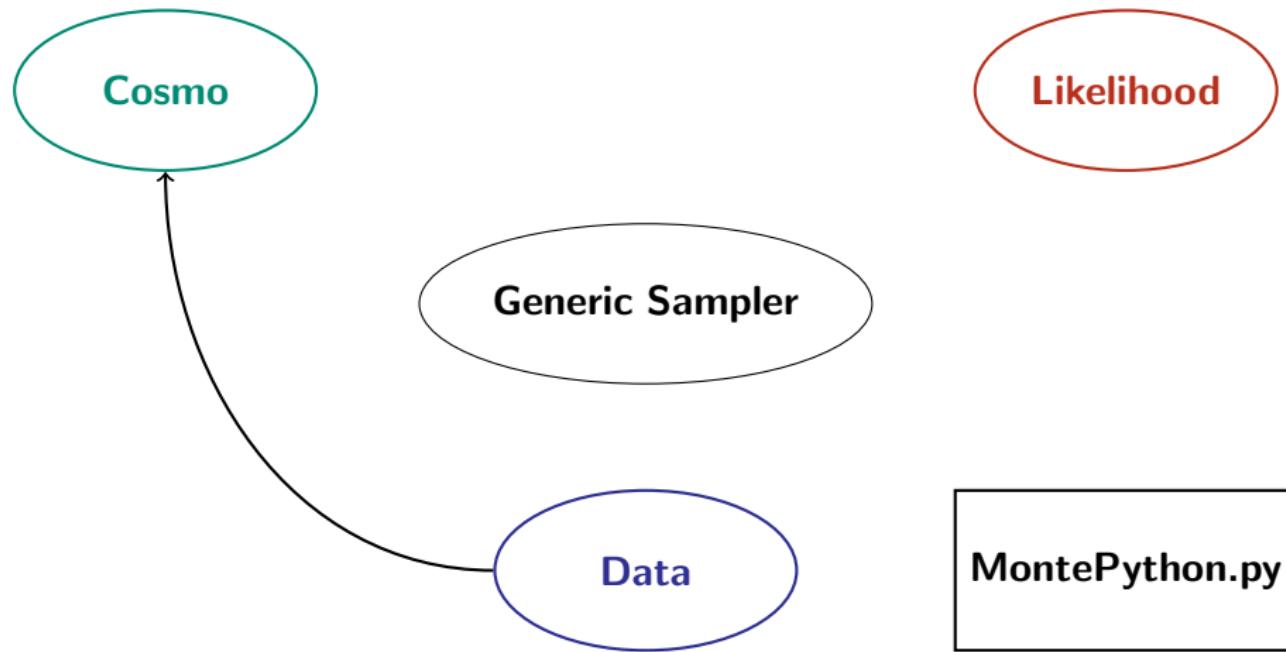
## Sampler

How to choose a new point?

- basic eigen-values/vector decomposition
- Cholesky decomposition (Planck) (-j fast)
- Nested Sampling with MultiNest (-m NS)
- Emcee with Cosmo Hammer (-m CH)

# Compute Likelihood

Set the cosmo



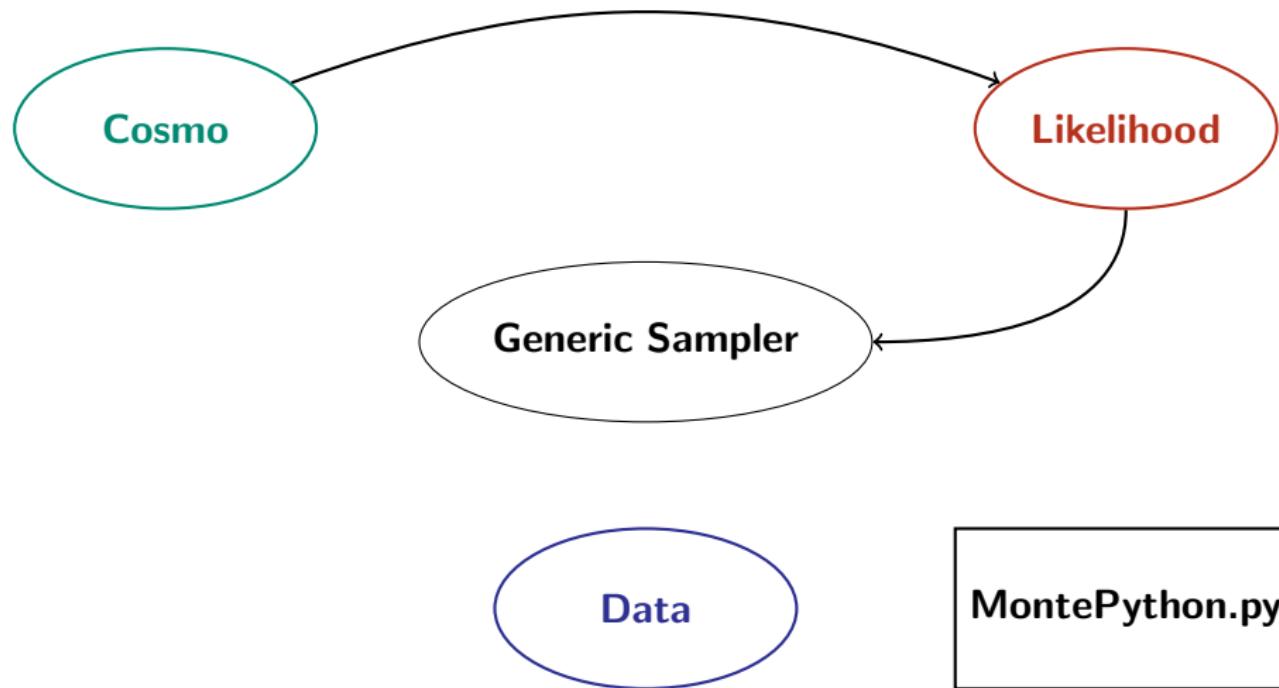
# Sampler III

## Compute likelihood

```
33 def compute_lkl(cosmo, data):  
  
370     if ((data.need_cosmo_update) or  
371         (not cosmo.state) or  
372         (data.jumping_factor == 0)):  
373  
374         # Prepare the cosmological module with the new set of  
            # parameters  
375         cosmo.set(data.cosmo_arguments)  
  
390     try:  
391         cosmo.compute()  
392     except CosmoComputationError:  
393         return data.boundary_loglike  
394     except CosmoSevereError, message:  
395         print str(message)  
396         raise io_mp.CosmologicalModuleError(  
397             "Something went wrong when calling CLASS")
```

# Compute Likelihood

For each likelihood



# Sampler III

## Compute likelihood

```
404     loglike = 0

410     for likelihood in data.lkl.itervalues():
411         if likelihood.need_update is True:
412             value = likelihood.loglkl(cosmo, data)
413             # Storing the result
414             likelihood.backup_value = value
415             # Otherwise, take the existing value
416         else:
417             value = likelihood.backup_value
418         loglike += value
```

...fiducial...

```
446     return loglike
```

# Sampler IV

Get the covariance matrix

## Main ideas

- stores values without scale factors for numerical reason

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Get the covariance matrix

## Main ideas

- stores values without scale factors for numerical reason
- automatic handling of parameters
- **computes eigen vectors, values, and Cholesky decomposition**

# Likelihood class

Heavily object oriented

in `likelihood_class.py` are defined:

- the basic `Likelihood` class (parent of all others)

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- `Likelihood_mpk` (WiggleZ, Euclid)

# Likelihoods

## Implementation

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# Likelihoods

## Implementation

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- `likelihoods/something/__init__.py` and
- `likelihoods/something/something.data`
- always `inherit` at least from: `Likelihood`

# Analyze

called with:

```
python montepython/MontePython.py info folder
```

# Analyze

## Convergence Computation

Gelman-Rubin Diagnostic:

variance between chains = variance within chains

This gives a number  $R$ , that must be  $< 0.01$ .

Beware of this number computed for a single file!

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pdf output (or png): triangle plot and 1-dimensional  
**marginalized posterior** and **Mean likelihood** (visual  
indication of convergence)

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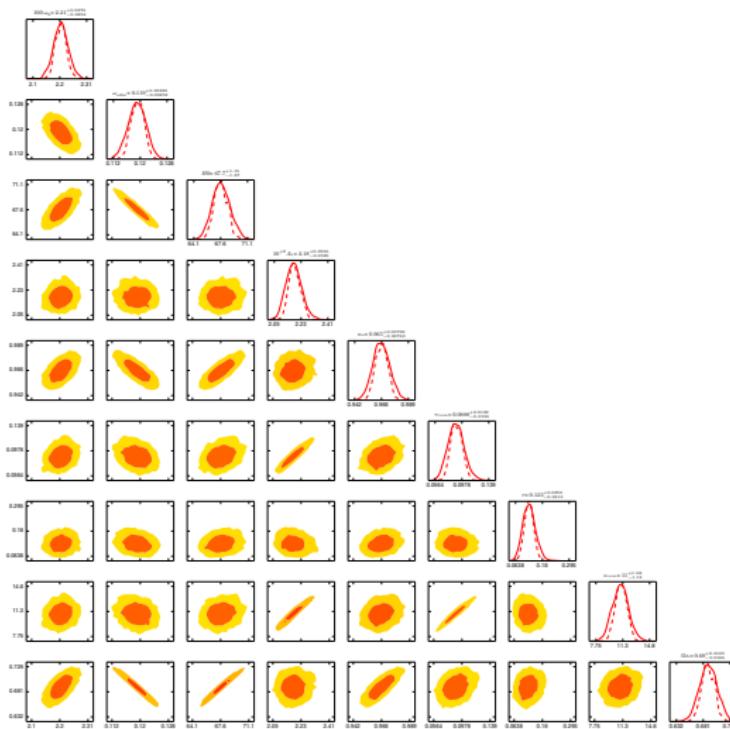
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## Output Files

.covmat, .v\_info, h\_info, .bestfit, .log

# Analyze



# Analyze

## horizontal info

```
param names   : Omega_L          h
R-1 values   : 0.000055  0.000015
Best Fit     : 7.918934e-01  7.339652e-01
mean         : 7.781465e-01  7.308724e-01
sigma        : 6.793778e-02  2.382595e-02

1-sigma -   : -5.799105e-02 -2.310380e-02
1-sigma +   : 7.788452e-02  2.454811e-02
2-sigma -   : -1.418414e-01 -4.808685e-02
2-sigma +   : 1.333294e-01  4.764994e-02
3-sigma -   : -2.379013e-01 -7.257030e-02
3-sigma +   : 1.779086e-01  6.790438e-02

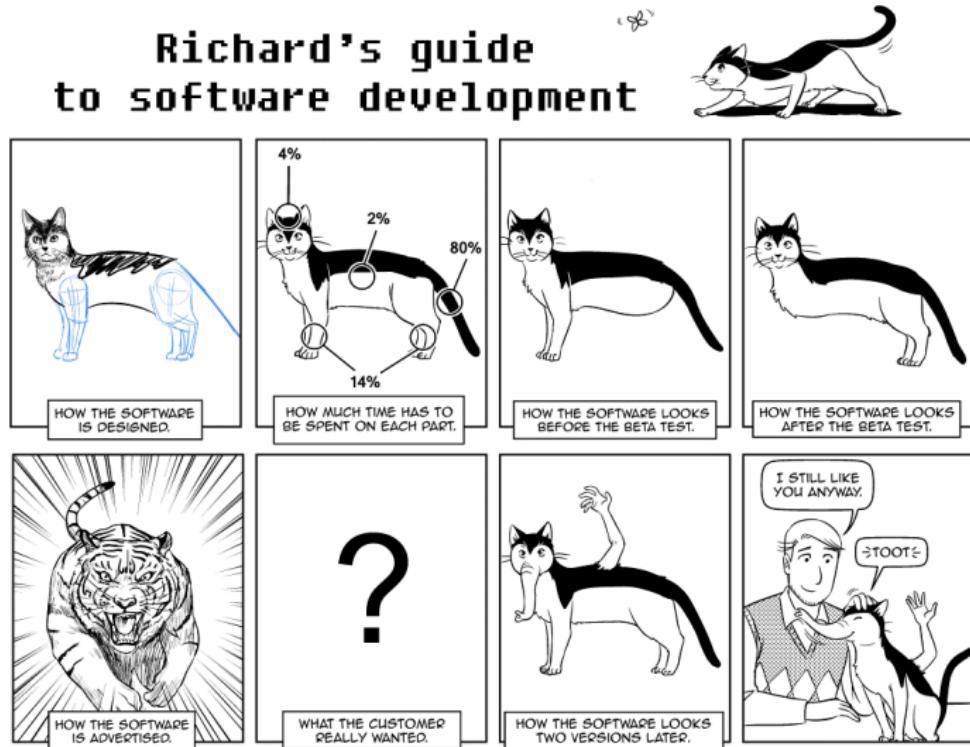
1-sigma >   : 7.201555e-01  7.077686e-01
1-sigma <   : 8.560310e-01  7.554205e-01
2-sigma >   : 6.363051e-01  6.827856e-01
2-sigma <   : 9.114759e-01  7.785224e-01
3-sigma >   : 5.402452e-01  6.583021e-01
3-sigma <   : 9.560552e-01  7.987768e-01
```

# Analyze

## other files

- run.log
- run.covmat
- run.tex

# Conclusion on Design



# Outline

1 Code Structure

2 Important Modules in (some) detail

3 Usage

- Complete work session example
- Analyzing and plotting the results

4 Practice with Monte Python

# Work session example

Test a model with running tilt with Planck data

- Copy `base.param` into `base_r.param`

```
data.experiments=['Planck_highl','Planck_lowl','lowlike']

data.parameters['omega_b'] = [2.2253, None, None, 0.028, 0.01, 'cosmo']
data.parameters['omega_cdm'] = [0.11919, None, None, 0.0027, 1, 'cosmo']
data.parameters['H0'] = [67.802, None, None, 1.2, 1, 'cosmo']
data.parameters['A_s'] = [2.2177, 0, None, 0.055, 1.e-9, 'cosmo']
data.parameters['n_s'] = [0.96229, 0, None, 0.0074, 1, 'cosmo']
data.parameters['tau_reio'] = [0.09463, 0, None, 0.013, 1, 'cosmo']
```

# Work session example

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- **Edit `base_r.param`, add  $r$ , impose self-consistency condition**

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```
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data.parameters['r'] = [0.00463, 0, None, 0.013, 1, 'cosmo']
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# Work session example

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- Copy `base.param` into `base_r.param`
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- **Launch a short run to see if it works:**

```
python montepython/MontePython.py -o chains/planck_r \
-p base_r.param -N 10
```

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- Edit `base_r.param`, add  $r$ , impose self-consistency condition
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`python montepython/MontePython.py -o chains/planck_r \`  
`-p base_r.param -N 10`
- **Launch many chains:** `export OMP_NUM_THREADS=2`  
`mpirun -np 24 python montepython/MontePython.py \`  
`-o chains/planck_r -N 10000`

# Analyzing and Plotting

## After running longer chains

- do: `python montepython/MontePython.py info chains/planck_r/*10000*`
- use the output covariance matrix as an input (new chains!)

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# Exercices

## I) Hst, SN, BAO

Make a  $\Lambda$ CDM run with Hubble Space Telescope, Super Novae and BAO data. Look at [montepython/likelihoods](#) for the names.

## II) Have fun with classy

Use the classy wrapper in a Python interpreter (notebook) to redo yesterday's exercices (see [https://github.com/lesgourg/class\\_public/wiki/Python-wrapper](https://github.com/lesgourg/class_public/wiki/Python-wrapper)).

## III) Installing Planck likelihood

and trying it with `base.param`

<http://pla.esac.esa.int/pla/aio/planckProducts.html>